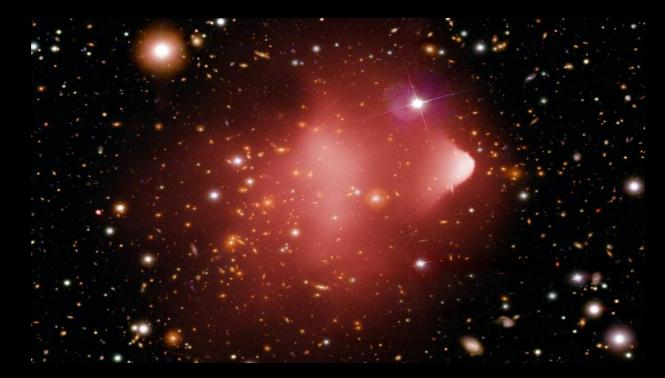
### Dark Matter and Dark Sectors at the LHC

Luca Lavezzo

On behalf of the Particle Physics Collaboration at MIT



### Unexplained phenomena

- Gravity
- Dark matter
- Dark energy
- Matter-antimatter asymmetry
- •••

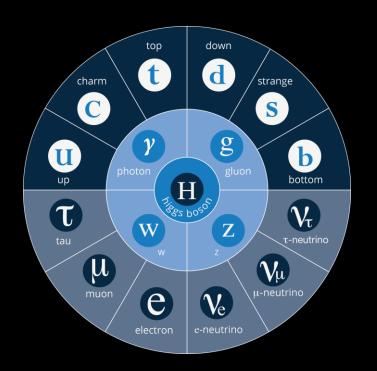
### Experimental tensions (?)

- $(g-2)_{\mu}$
- $m_W$
- $R(D^*)$
- *X*17
- •••

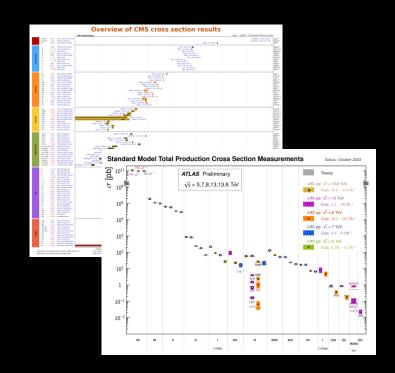
...

### Fine-tuning problems

- $\theta_{CP} \approx 0$
- Hierarchy problem
- Neutrino masses
- Choice of parameters



# Most consistent and precise theory in human history



For the first time, no clear indication about what the missing pieces are

### Unexplained phenomena

- Gravity
- Dark matter
- Dark energy
- Matter-antimatter asymmetry

•••

### Experimental tensions (?)

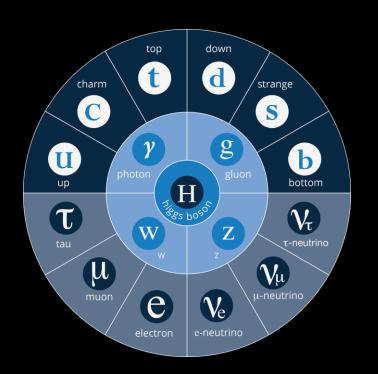
- $(g-2)_{\mu}$
- $m_W$  (CMS: nothing to see here!)
- $R(D^*)$
- *X*17

•••

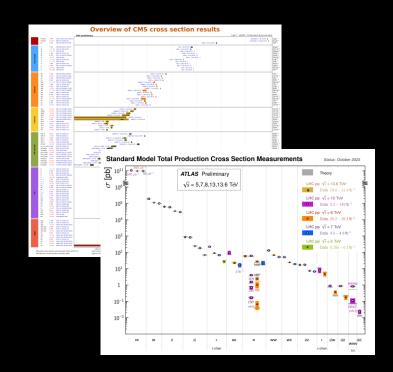
...

### Fine-tuning problems

- $\theta_{CP} \approx 0$
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# Most consistent and precise theory in human history



For the first time, no clear indication about what the missing pieces are Dark sectors (DS) can address any of these problems

- New interactions with the standard model (SM) can provide dark matter (DM) candidates
- New symmetries can solve other theoretical and finetuning problems
- New particles can explain experimental tensions
- Dark sectors have their own **dark charges**, so are stable under their conservation laws, and can have rich structure

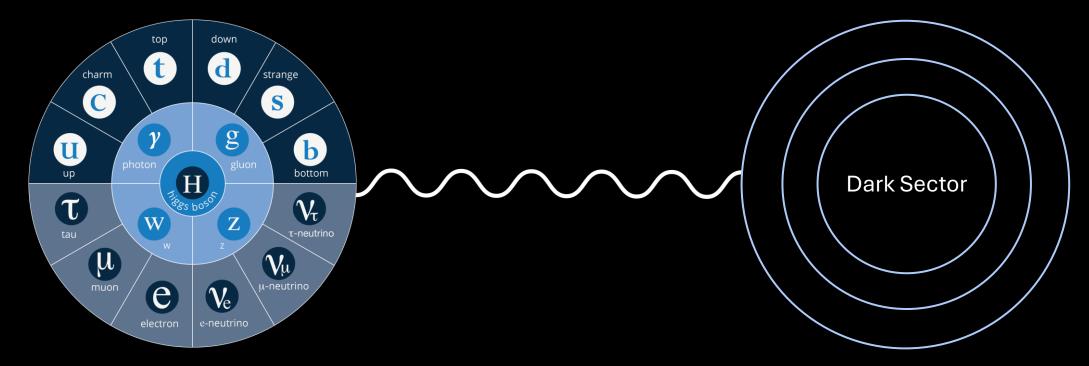
No hints about any details of the DS!

So, how do we start looking?

Dark Sector	

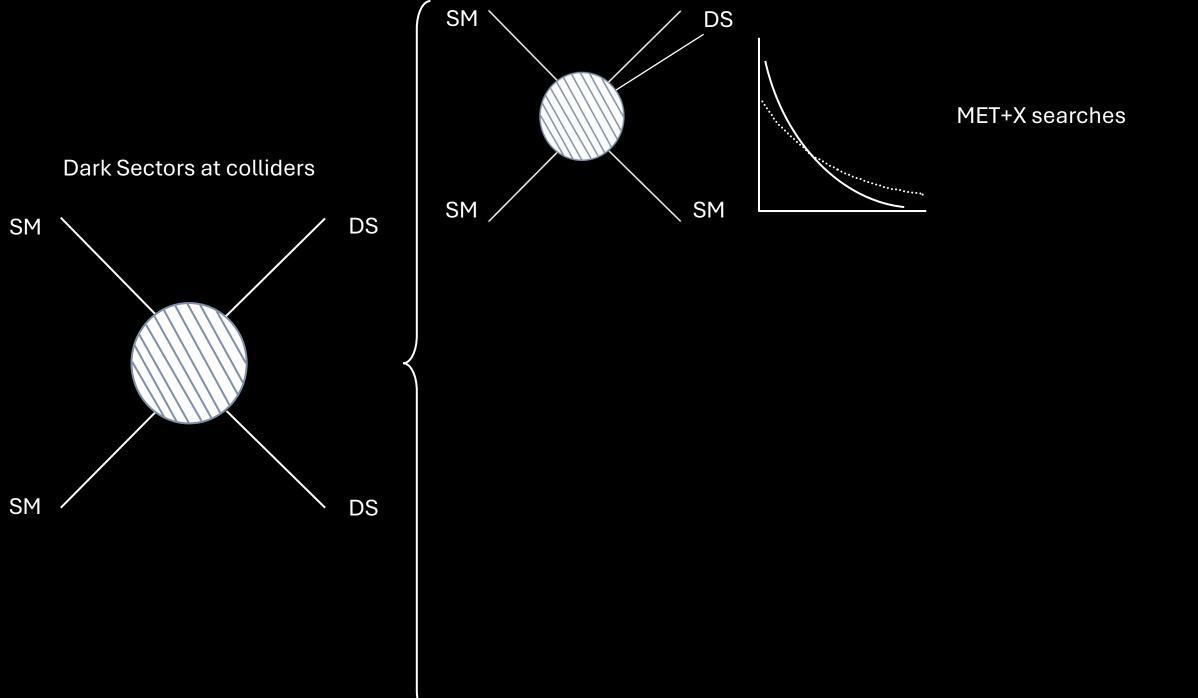
If there is any interaction between DS and SM, there needs to be a **portal** between the two.

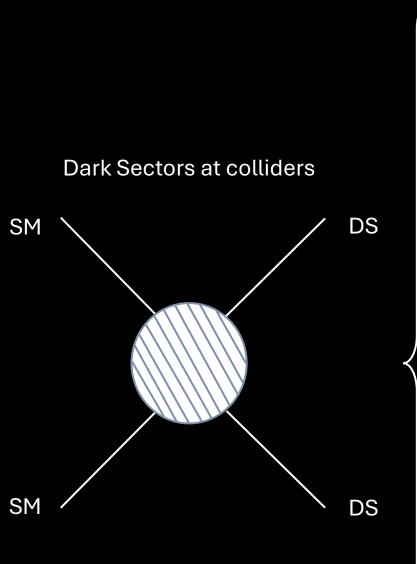
To avoid breaking SM symmetries, four commonly studied ways to communicate with DS: Spin-1 Portal, Spin-0 Portal, Fermion Portal, Neutrino Portal

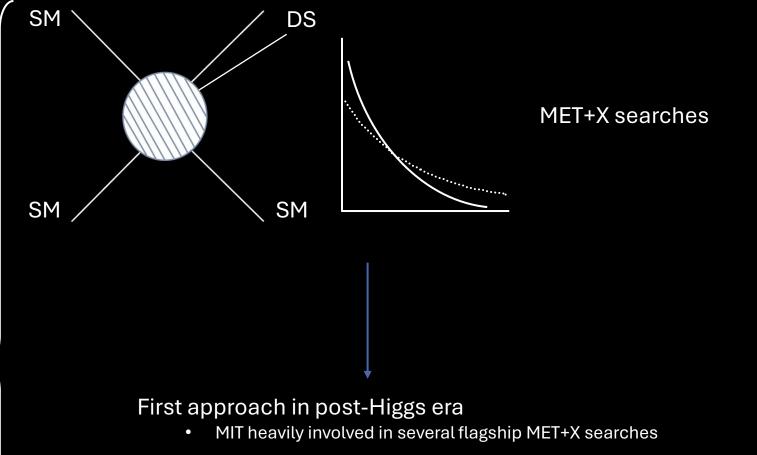


Many portals within direct reach, others can show up at lower energies due to quantum mechanical mixing

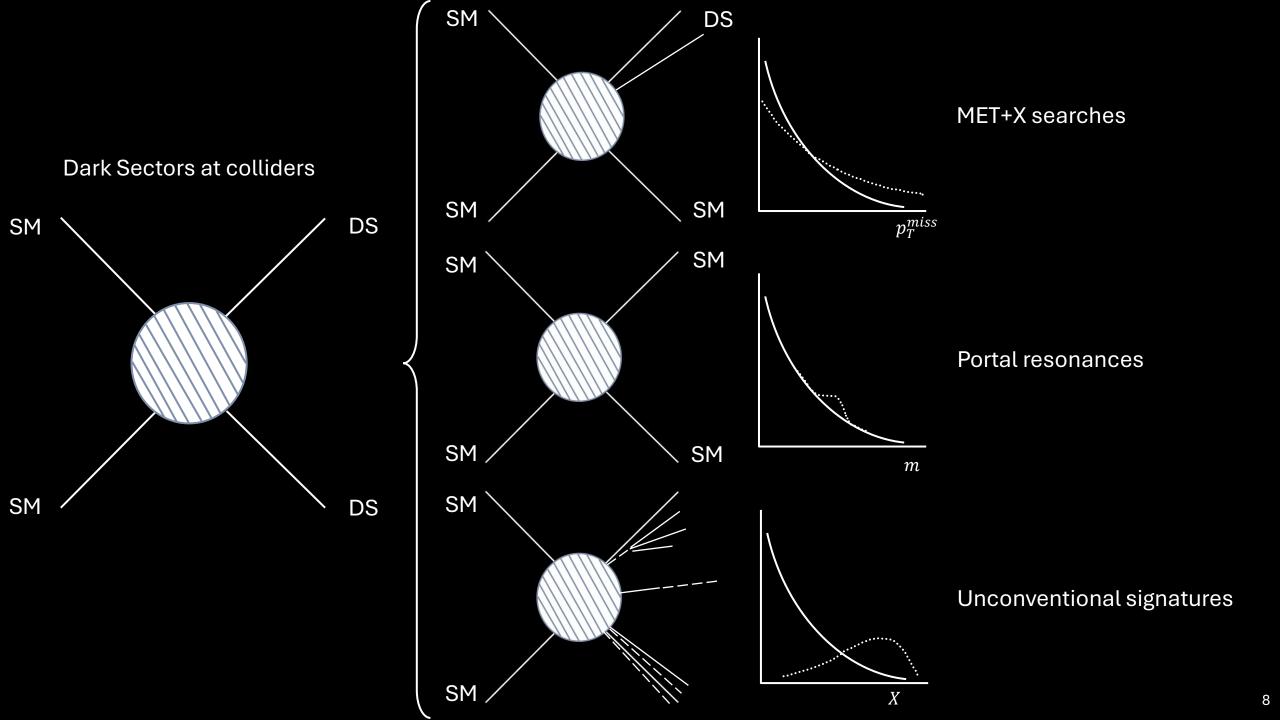
How can we probe this at colliders?

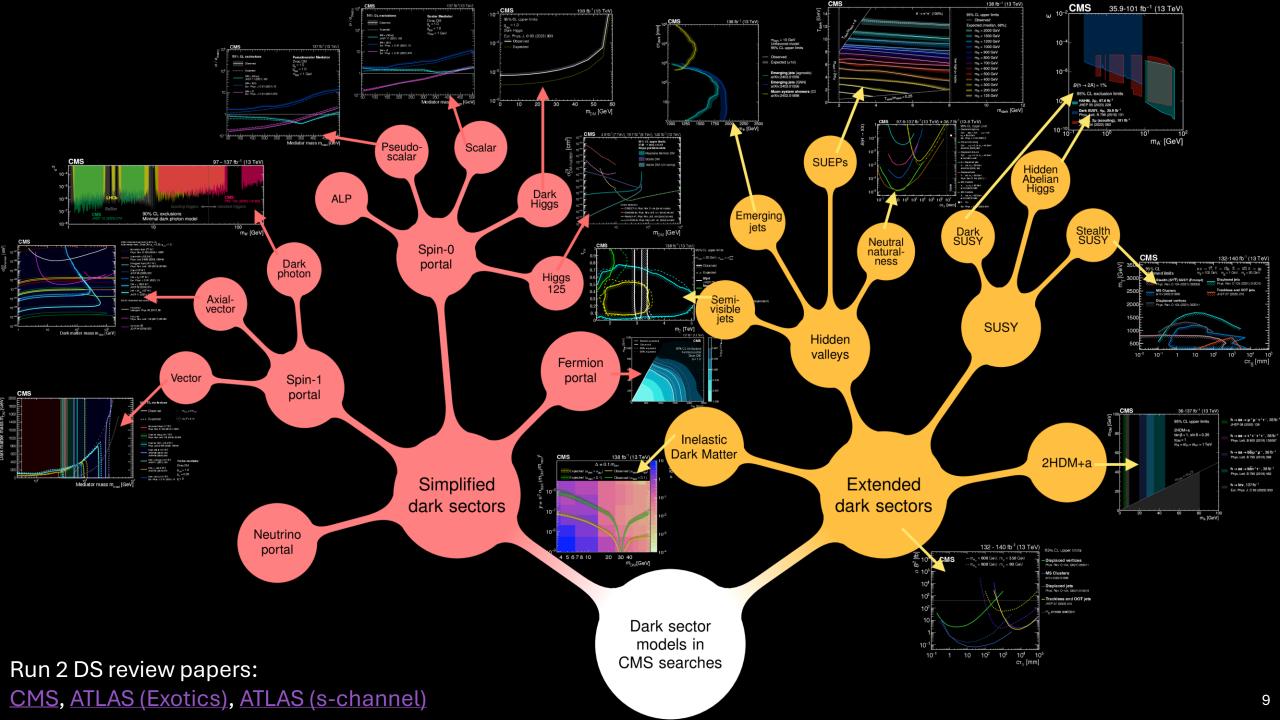


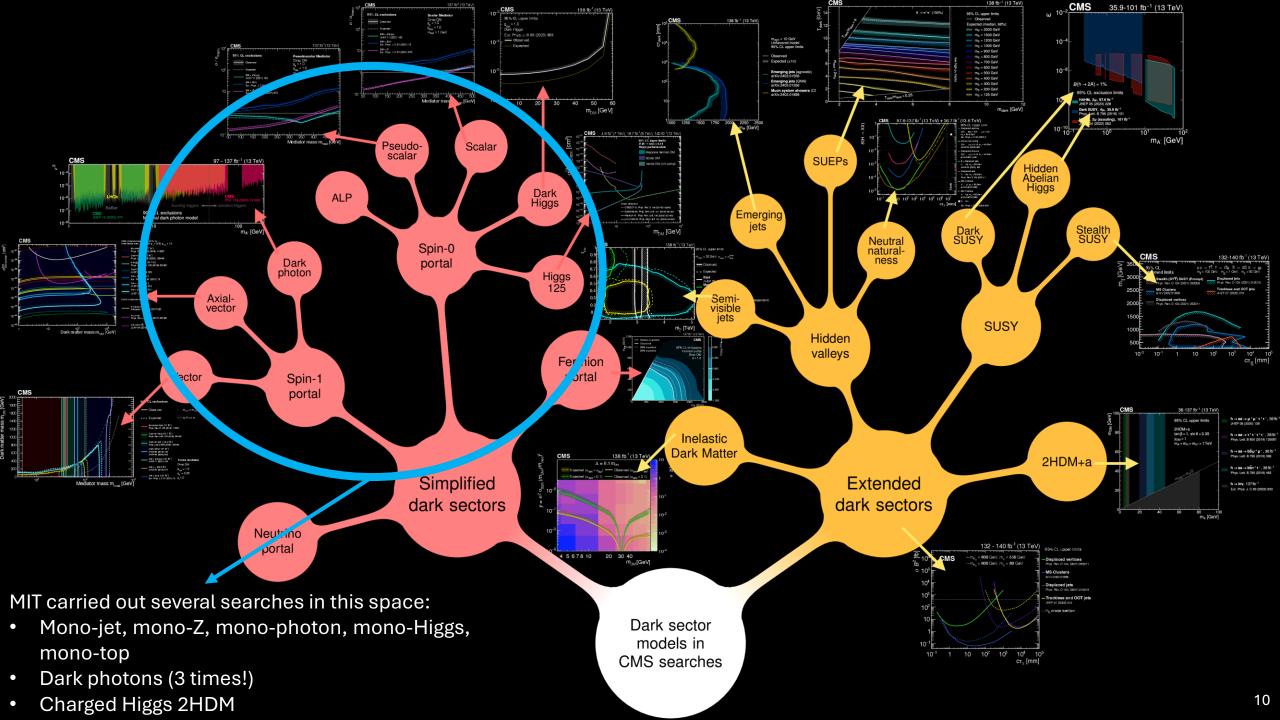


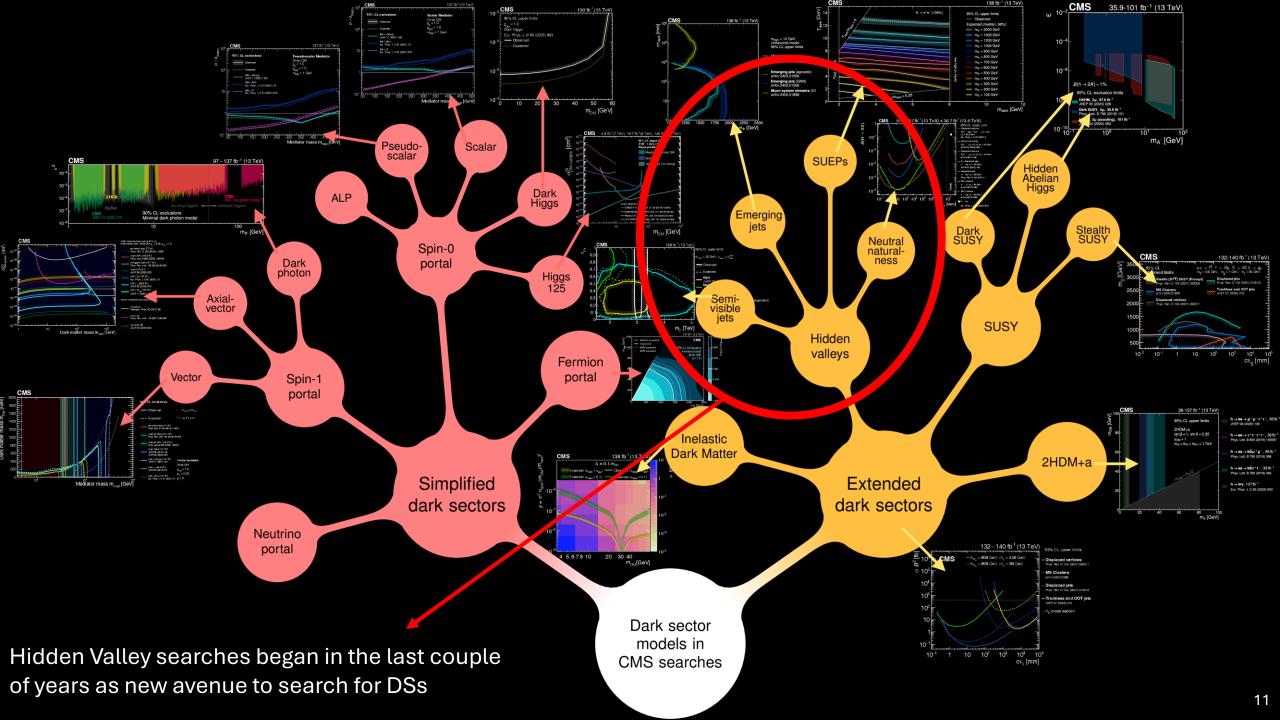


Series of null results that heavily constrained BSM theories  $\rightarrow$  need to look elsewhere









### Hidden Valleys

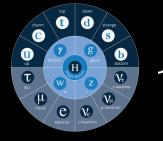
Standard Model Hidden Valley



 $\sim \sim SU(N_{dark \ color})$ 

# Why Hidden Valleys?

Standard Model Hidden Valley



 $\sim \sim SU(N_{dark \ color})$ 

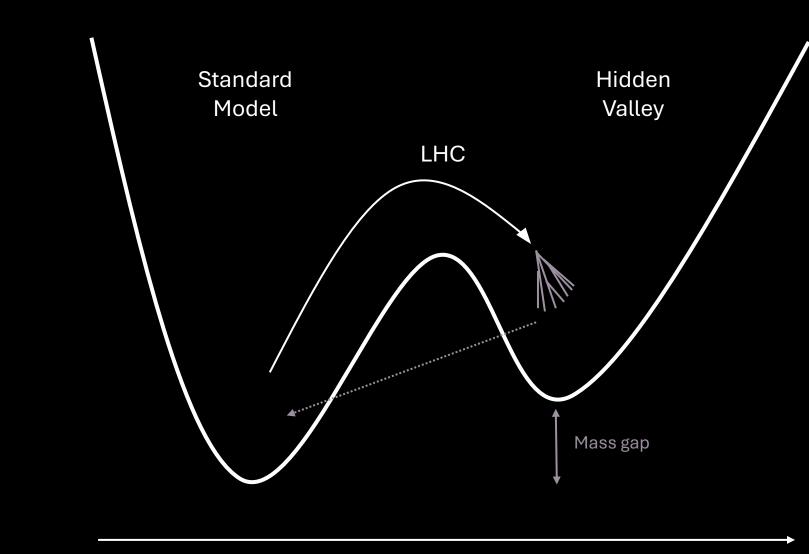


Rich phenomenology at colliders!

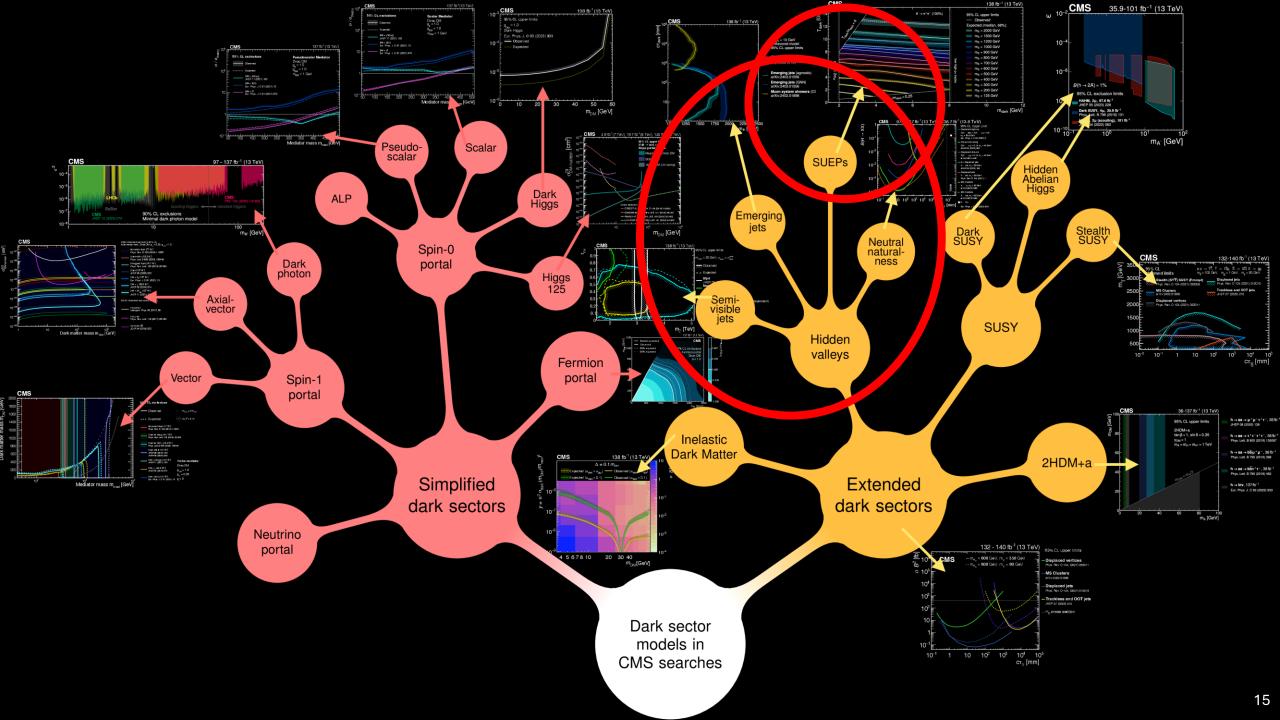
Many new possible signatures that would've been missed in previous searches: semi-visible jets, emerging jets, SUEPs. Strong, confining (QCD-like) force can give rise to composite dark hadrons which fulfill DM criteria (stable, small interactions, heavy)



### How to find Hidden Valleys at colliders



Energy

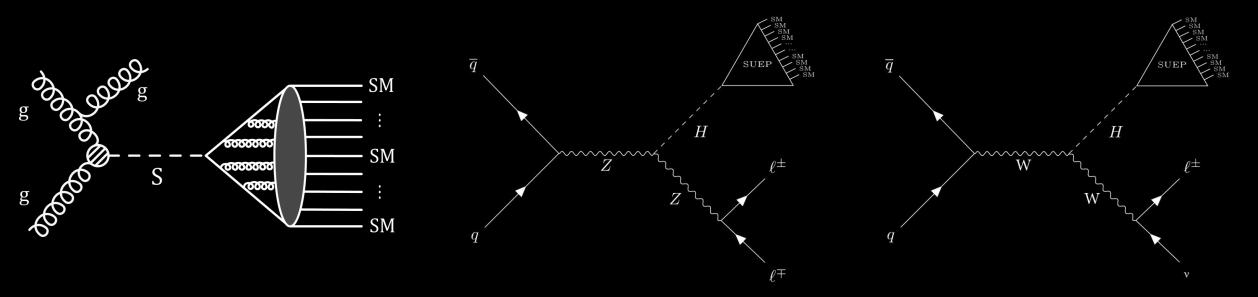


## The SUEP program at CMS

### Five ongoing searches for SUEPs using Run 2 data at CMS

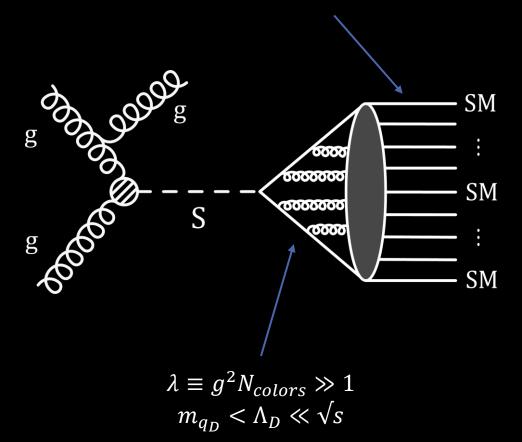
MIT leading the effort along with BU, Fermilab

- Began the effort in 2021, led the group since
- Developed a common framework to perform these analyses
- Gluon fusion channel, led by MIT, just <u>accepted to PRL</u>, with Editor's Highlight



### A Hidden Valley paradigm: Soft Unclustered Energy Patterns (SUEPs)

Promptly decay back to SM

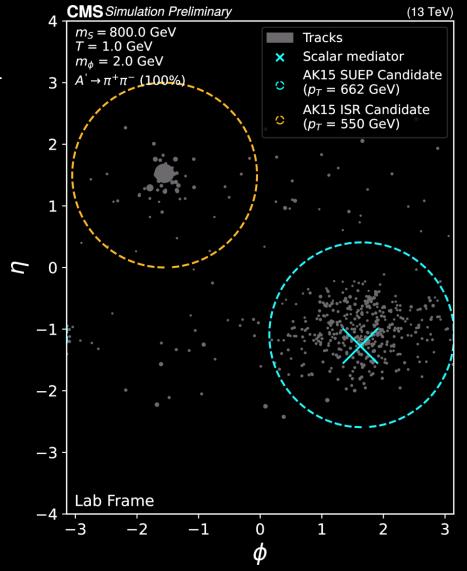


Efficient, isotropic showering window

### A 'real-life' SUEP at the LHC

QCD

Low multiplicity of higher momentum particles, produced collinearly

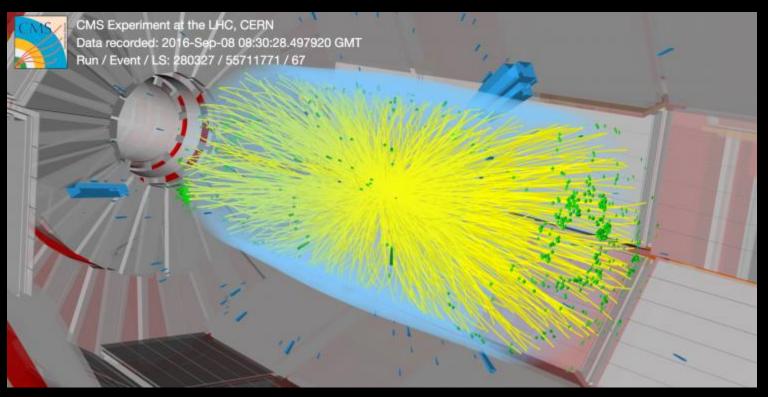


SUEP

High multiplicity of low momentum particles, isotropic in its own frame

## Experimental challenge

### "A spray of low $p_T$ tracks? Have you heard of pileup?"

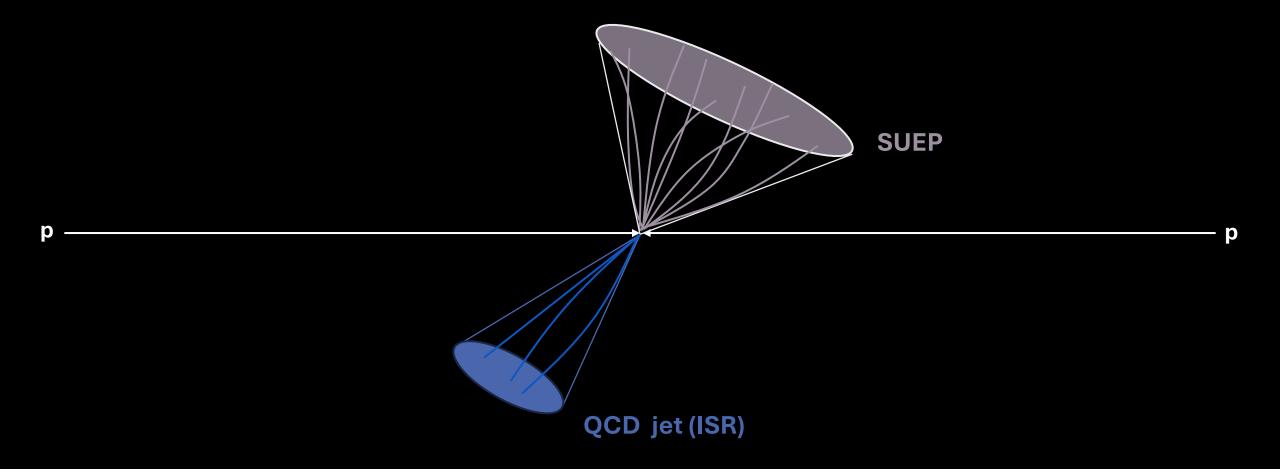


Average event at CMS: quite hard to distinguish SUEP from so-called "pileup", the additional overlapping pp interactions that happen concurrently with the interaction of interest.

From https://cms.cern/news/how-cms-weeds-out-particles-pile

### The gluon fusion channel

Use events with high hadronic to select recoiling SUEP-ISR system → Can now trigger on these events using the hadronic activity

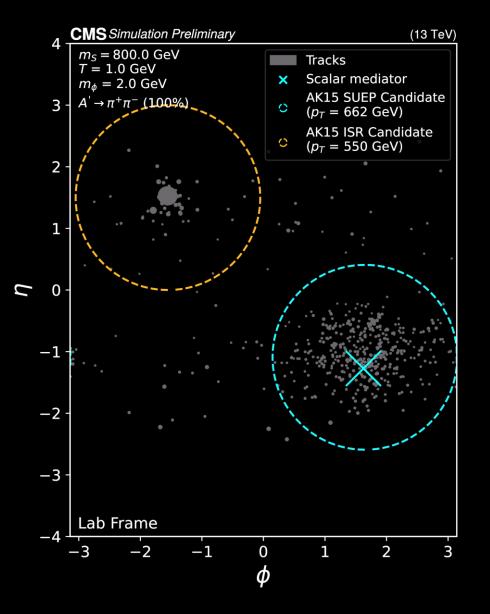


### Reconstruction

- SUEPs are isotropic, soft sprays of particles
- → Evade typical reconstructed objects: even typical jets can't capture the full sprays that the signals produce

Solution:

- Cluster 'wide-jets' using anti-kt algorithm with large radii
- Novel type of reconstructed object, makes these analyses possible
- Displays CMS' ability to push beyond its original design

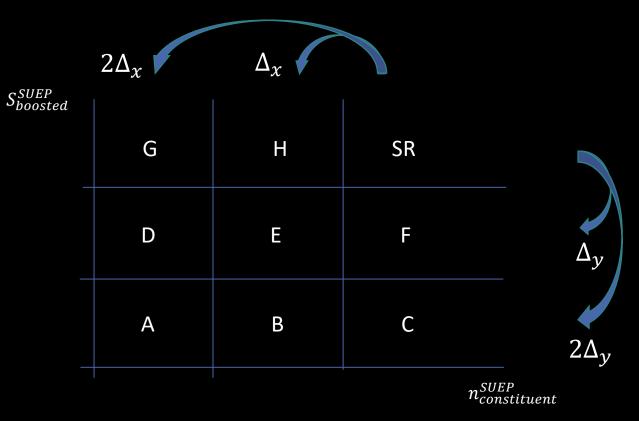


# **Background Estimation**

- Premiered a novel data-driven estimation technique (arxiv:1906.10831)
  - Fully data-driven method to predict background in signal region (SR)
  - Account for linear correlations in variables
  - Shape prediction for SR

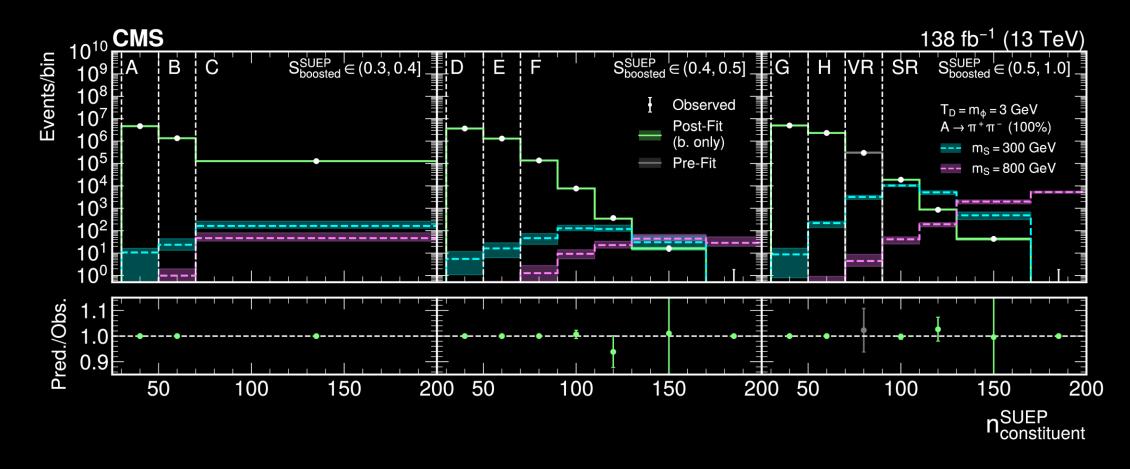
$$SR^{Bin i} \approx F^{Bin i} \frac{H^2 F D^2 B^2}{G C A E^4} + O(\Delta^4)$$

Scaling factor applied to F histogram



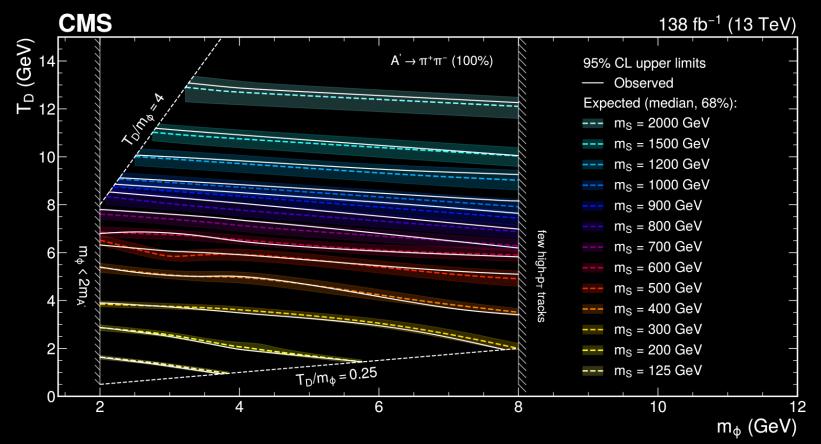
• Now used on other analyses, including mW!

### Results



- Huge number of events and extended ABCD method yield a very precise prediction of the notoriously difficult QCD jet processes
  - Systematics on the ABCD prediction constrained through the control regions (A-H) were key to this

# Limits on SUEPs

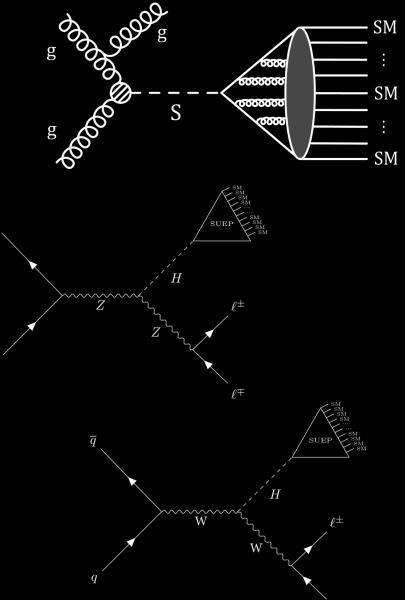


- First limits on SUEPs
  - Result can be re-interpreted in other BSM models that produce spherical, numerous jets
  - Similar analysis strategy might target instantons one day
  - Add to searches for emerging jets and semi-visible jets in the effort to search for Hidden Valleys
- Accepted by PRL with Editor's Highlight

## **Other channels**

MIT involved in other channels to push the SUEP program

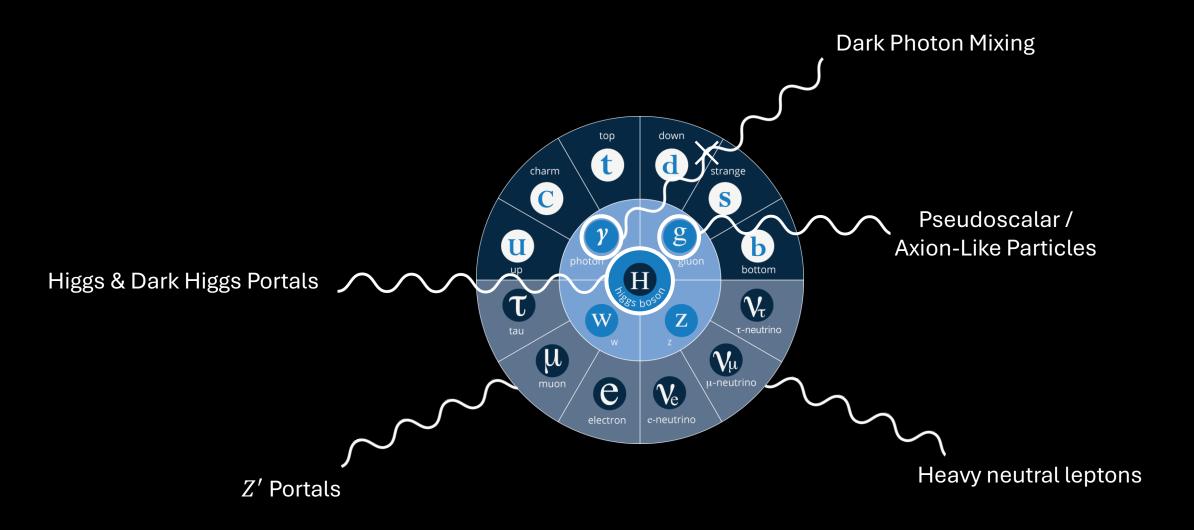
- Gluon fusion in scouting:
  - Make use of 'scouting' dataset
  - Relatively unused still, but very interesting: huge yields of less-detailed data
  - Can target much lower mass scalar mediators that suffer from low rates in the standard 'offline' analysis
- Associated production with a vector boson (W and Z):
  - Target production of a Higgs (or Higgs-like) scalar that decays to SUEPs
  - Providing the analysis in a model-independent framework that can be re-interpreted by theorists



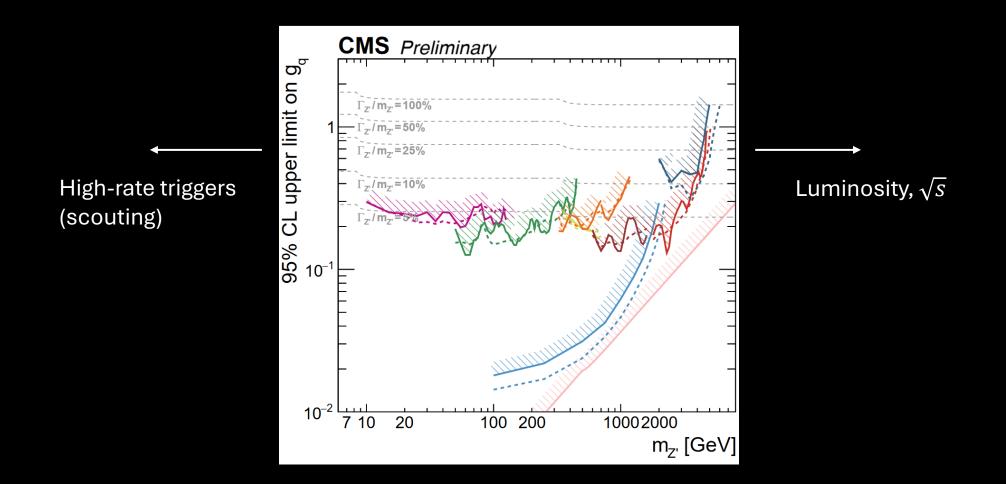
## Conclusions

- Searches for dark sectors are the new frontier for BSM at colliders
- Hidden Valleys provide a framework for new dark sectors that are both interesting for theory and experiments
- MIT had led the way on SUEPs, an important Hidden Valley paradigm, providing new searches that have been of great interest to the community
  - New signatures
  - New reconstructions
  - New statistical techniques
- Future:
  - Many other channels ongoing
  - Feedback from theorists: modified SUEPs with different assumptions, decays to SM, etc.
  - HL-LHC can really target SUEP models via triggering on tracks

# Backup



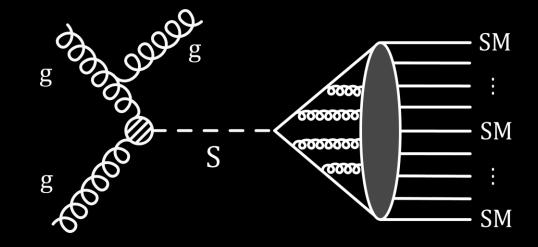
### Improving MET+X and Resonance Searches

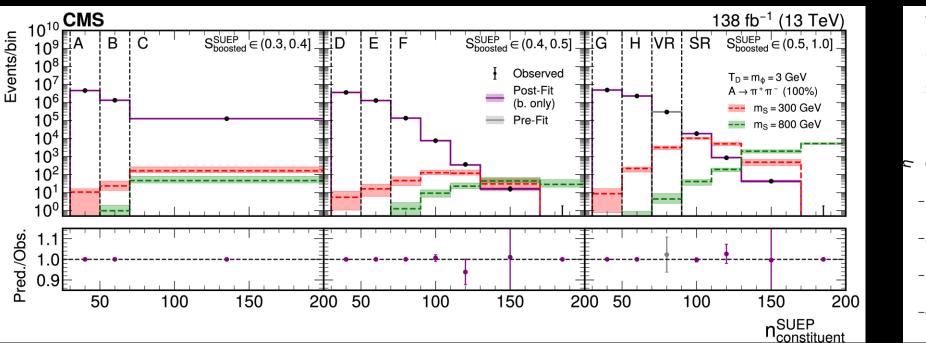


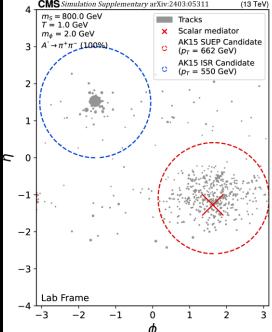
+ better methods: event selection, jet tagging, background estimations, etc.

# Soft Unclustered Energy Patterns

- Strongly coupled dark sector connected via scalar portal
- Large 't Hooft coupling in quasi conformal dark sector
  - Long, efficient showering window, which produces spherical, high multiplicity jets
- Trigger on events with SUEP recoiling against ISR
- Background prediction using extended ABCD method

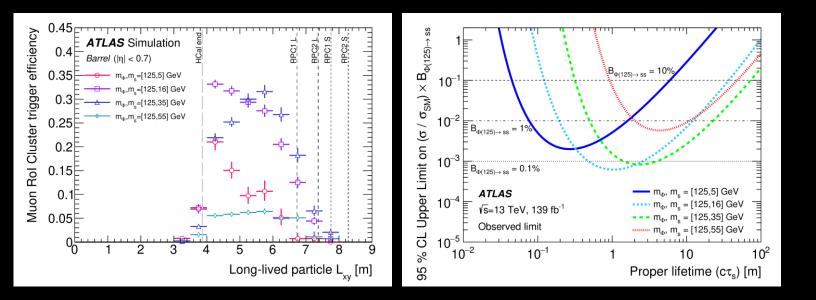


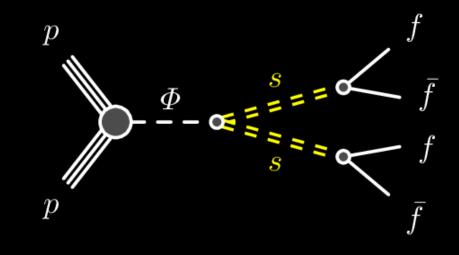




### Unconventional Signatures: Displaced Jets in $\mu$ Spectrometer

- Dark particles *s* produced, travel ~m, decay to *ff*, which shower
  - Showers in the  $\mu$  spectrometer
- Dedicated trigger to look for several tracks in the  $\mu$  spectrometer within a jet's typical radius,  $\Delta R = 1.5$
- Dedicated reconstruction for displaced decays in the spectrometer

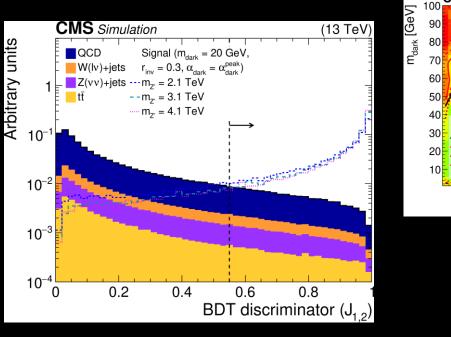


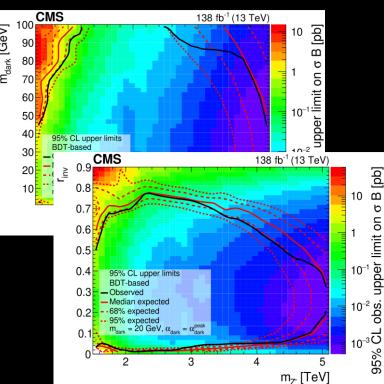


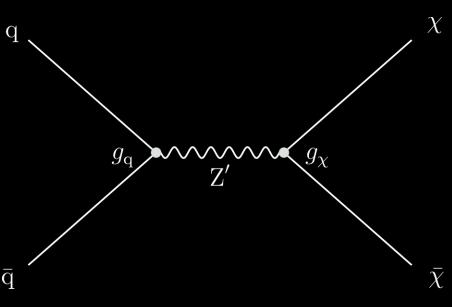
- Set limits on  $c\tau, \sigma$
- Assumptions on DS structure, BRs, masses, etc. for these limits

## Unconventional Signatures: Semivisible Jets

- $\chi$  produced through a Z' mediator, which shower, some decaying back to SM quarks, some staying in DS, controlled by ratio  $r_{inv}$ 
  - Collimated mixtures of visible and invisible particles
- Use  $p_T^{miss}$ , N-subjettiness, energy correlators, soft-drop mass in a BDT to discriminate between SVJ and SM jets



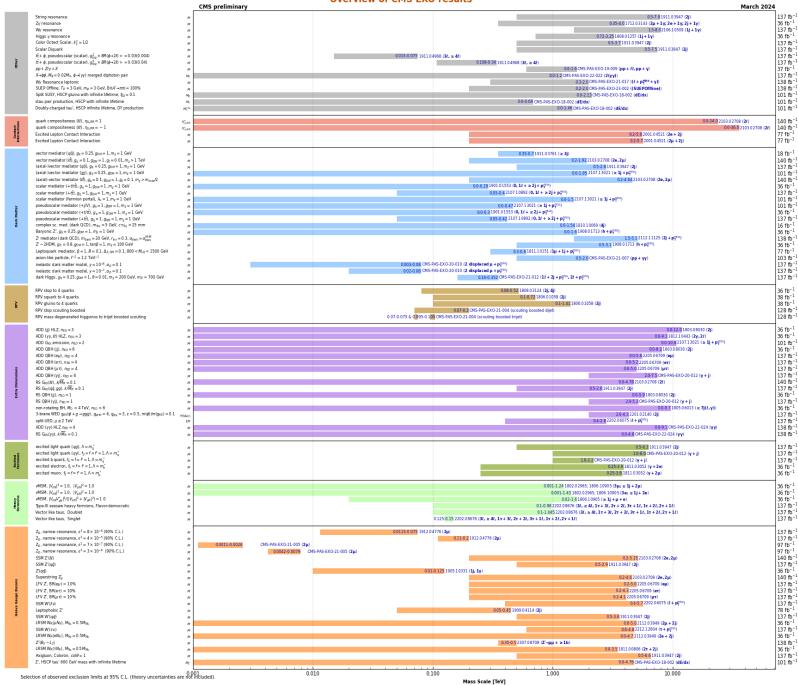




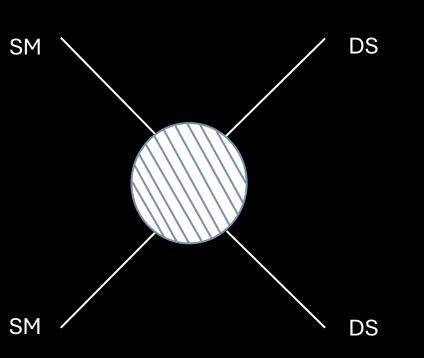
- Set limits on  $m_{Z'}, m_{dark}, r_{inv}$
- Assumptions on dark sector, such as showering, gauge structure, etc.

#### **EXO-CMS Summary Plots**

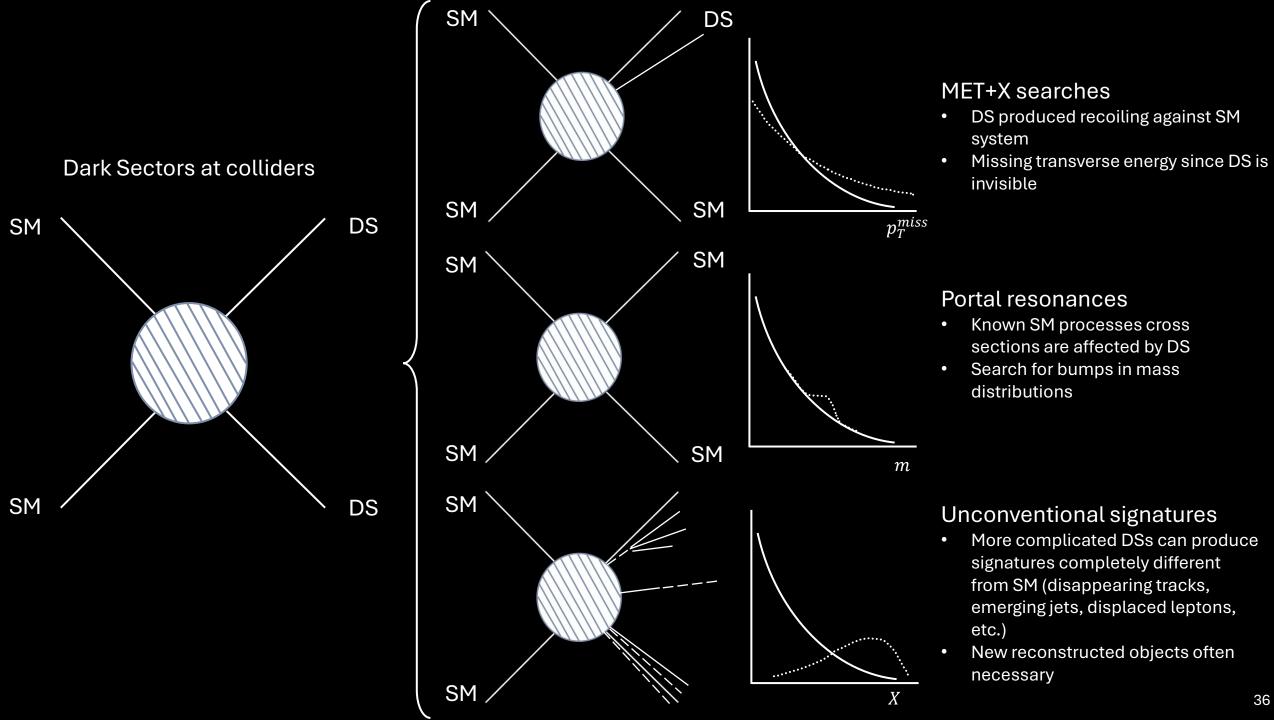
#### Overview of CMS EXO results

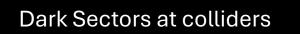


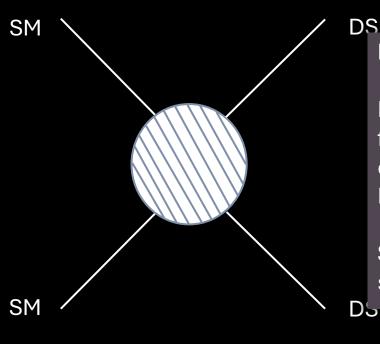
ATLAS Heavy P Status: March 2023	article Searches* -	95% CL Upper	Exclusion Limits	$\int \mathcal{L} dt = (3.6 - 139) \text{ fb}^{-1}$	<b>LAS</b> Preliminary $\sqrt{s} = 13 \text{ TeV}$	ATLAS Long Status: March 2023		e Searches	s* - 95% CL Exclusio		<b>ATL</b> t = (32.8 - 139) fb <sup>-1</sup>	<b>AS</b> Preliminary $\sqrt{s} = 13 \text{ TeV}$
Model	ℓ,γ Jets† E <sub>T</sub> <sup>miss</sup> ∫⊥dt	[fb <sup>-1</sup> ]	Limit	0	Reference	Model	Signature f	$\mathcal{L} dt [fb^{-1}]$	Lifetime limit	0		Reference
ADD $G_{KK} + g/q$ ADD non-resonant $\gamma\gamma$ ADD OBH ADD BH multijet RS1 $G_{KK} \rightarrow \gamma\gamma$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ms Mth Mth	45	11.2 TeV n = 2 8.6 TeV n = 3 HLZ NLO 9.4 TeV n = 6, M <sub>D</sub> = 3 TeV, rot BH k/M <sub>FF</sub> = 0.1	2102.10874 1707.04147 1910.08447 1512.02586 2102.13405	RPV $\tilde{t} \rightarrow \mu q$ RPV $\tilde{\chi}_1^0 \rightarrow eev/e\mu v/\mu \mu$ RPV $\tilde{\chi}_1^0 \rightarrow qqq$	displaced vtx + muon v displaced lepton pair displaced vtx + jets	136 $\tilde{t}$ lifetime 32.8 $\tilde{\chi}_{1}^{0}$ lifetime 139 $\tilde{\chi}_{2}^{0}$ lifetime		0.003-6.0 m 0.003-1.0 m 0.00135-9.0 m	$m(\tilde{t}) = 1.4$ TeV $m(\tilde{q}) = 1.6$ TeV, $m(\tilde{\chi}_1^0) = 1.3$ TeV $m(\tilde{\chi}_1^0) = 1.0$ TeV	2003.11956 1907.10037 2301.13866
Bulk RS $G_{KK} \rightarrow WW/ZZ$ Bulk RS $g_{KK} \rightarrow tt$	multi-channel 36.1 1 e, μ ≥1 b, ≥1J/2j Yes 36.1	G <sub>KK</sub> mass g <sub>KK</sub> mass	2.3 TeV 3.8 TeV	$k/\overline{M}_{Pf} = 1.0$ $\Gamma/m = 15\%$	1808.02380 1804.10823	$\operatorname{GGM} \tilde{\chi}_1^0 \to Z \tilde{G}$	displaced dimuon	32.9 $\tilde{\chi}_1^0$ lifetime		0.029-18.0		1808.03057
$2UED / RPP$ SSM Z' $\rightarrow \ell\ell$	1 e, μ ≥2 b, ≥3 j Yes 36.1 2 e, μ 139	KK mass Z' mass	1.8 TeV	Tier (1,1), $\mathcal{B}(A^{(1,1)} \rightarrow tt) = 1$ 1 TeV	1803.09678	GMSB	non-pointing or delayed $\gamma$	139 $\tilde{\chi}_1^0$ lifetime		0.24-2.4 m	$m(\tilde{\chi}_1^0,\tilde{G}){=}60,20~{\rm GeV},\mathcal{B}_{\mathcal{H}}{=}2\%$	2209.01029
SSM $Z' \rightarrow \tau \tau$ Leptophobic $Z' \rightarrow bb$	2τ – – 36.1 – 2b – 36.1	Z' mass Z' mass	2.42 TeV 2.1 TeV		1709.07242 1805.09299	$GMSB \tilde{\ell} \rightarrow \ell \tilde{G}$	displaced lepton	139 <i>i</i> lifetime		6-750 mm	$m(\tilde{\ell}) = 600 \; { m GeV}$	2011.07812
Leptophobic $Z' \rightarrow tt$ SSM $W' \rightarrow \ell y$	0 e, μ ≥1 b, ≥2 J Yes 139 1 e, μ − Yes 139	W' mass		6.0 TeV	2005.05138 1906.05609	GMSB $\tilde{\tau} \rightarrow \tau \tilde{G}$	displaced lepton	139 7 lifetime	9	270 mm	$m(\tilde{\ell})=200~{ m GeV}$	2011.07812
SSM $W' \rightarrow \tau v$ SSM $W' \rightarrow tb$	1 τ − Yes 139 − ≥1 b, ≥1 J − 139	W' mass	4.4 1		ATLAS-CONF-2021-025 ATLAS-CONF-2021-043	AMSB $pp \rightarrow \tilde{\chi}_1^+ \tilde{\chi}_1^0, \tilde{\chi}_1^+ \tilde{\chi}_1$	-	136 $\tilde{\chi}_1^{\pm}$ lifetime		0.06-3.06 m	$m( ilde{\chi}_1^{\pm})=$ 650 GeV	2201.02472
$ \begin{array}{c} \overleftarrow{\mathbf{w}} \\ \mathbf{g} \\ \mathbf{g} \\ \mathbf{H} \mathbf{V} \mathbf{T} & \mathbf{W}' \rightarrow WZ \text{ model } \mathbf{B} \\ \mathbf{H} \mathbf{V} \mathbf{T} & \mathbf{W}' \rightarrow WZ \rightarrow \ell \nu \ \ell' \ell' \text{ model } \mathbf{M} \end{array} $		W' mass 340	) GeV	$g_V c_H = 1, g_f = 0$	2004.14636 2207.03925	AMSB $pp \rightarrow \tilde{\chi}_1^* \tilde{\chi}_1^0, \tilde{\chi}_1^+ \tilde{\chi}_1^0$		139 $\tilde{\chi}_1^{\pm}$ lifetime		0.3-30.0 m	$m(\tilde{\chi}_1^{\pm})=600 \text{ GeV}$	2205.06013
HVT $Z' \rightarrow WW$ model B LRSM $W_R \rightarrow \mu N_R$	1 <i>e</i> ,μ 2j/1J Yes 139 2μ 1J - 80		3.9 Te\ 5.0	V $g_V = 3$ O TeV $m(N_R) = 0.5$ TeV, $g_L = g_R$	2004.14636 1904.12679	Stealth SUSY		36.1 Š lifetime		0.1-519 m > 0.45 m	$\mathcal{B}(\tilde{g} \to \tilde{S}g) = 0.1, \ m(\tilde{g}) = 500 \text{ Ge}$	
Cl qqqq Cl ℓℓqq	- 2 j - 37.0 2 e, μ 139	Λ Λ		21.8 TeV η <sub>LL</sub> 35.8 TeV η	1703.09127 2006.12946	Split SUSY	large pixel dE/dx displaced vtx + E <sub>T</sub> <sup>miss</sup>	139 g lifetime 32.8 g lifetime		> 0.45 m 0.03-13.2 m	$m(\tilde{g}) = 1.8 \text{ TeV}, m(\tilde{\chi}_1^0) = 100 \text{ GeV}$ $m(\tilde{g}) = 1.8 \text{ TeV}, m(\tilde{\chi}_1^0) = 100 \text{ GeV}$	2205.06013 1710.04901
Cl eebs Cl µµbs Cl tttt	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	٨	1.8 TeV 2.0 TeV 2.57 TeV	$egin{array}{llllllllllllllllllllllllllllllllllll$	2105.13847 2105.13847 1811.02305	Split SUSY		36.1 g lifetime		0.0-2.1 m	$m(\tilde{g}) = 1.8 \text{ TeV}, \ m(\tilde{\chi}_1) = 100 \text{ GeV}$ $m(\tilde{g}) = 1.8 \text{ TeV}, \ m(\tilde{\chi}_1^0) = 100 \text{ GeV}$	
Axial-vector med. (Dirac DM) Pseudo-scalar med. (Dirac DM) Vector med. Z'-2HDM (Dirac Pseudo-scalar med. 2HDM+a	DM) 0 e, µ 2 b Yes 139	m <sub>med</sub> 3 m <sub>Z'</sub>	76 GeV 3.8 TeV 3.0 TeV 800 GeV	$\begin{array}{c} g_{q}=0.25, \ g_{\chi}=1, \ m(\chi)=10 \ {\rm TeV} \\ g_{q}=1, \ g_{\chi}=1, \ m(\chi)=1 \ {\rm GeV} \\ {\rm tan} \ \beta=1, \ g_{Z}=0.8, \ m(\chi)=100 \ {\rm GeV} \\ {\rm tan} \ \beta=1, \ g_{\chi}=1, \ m(\chi)=10 \ {\rm GeV} \end{array}$	ATL-PHYS-PUB-2022-036 2102.10874 V 2108.13391 ATLAS-CONF-2021-036	$H \to ss$ $H \to ss$	2 MS vertices 2 low-EMF trackless jets	139s lifetime139s lifetime		0.31-72.4 m 0.19-6.94 m	m(s)= 35 GeV m(s)= 35 GeV	2203.00587 2203.01009
Scalar LQ 1 <sup>st</sup> gen Scalar LQ 2 <sup>nd</sup> gen Scalar LQ 3 <sup>rd</sup> gen Scalar LQ 3 <sup>rd</sup> gen Scalar LQ 3 <sup>rd</sup> gen Scalar LQ 3 <sup>rd</sup> gen Vector LQ 3 <sup>rd</sup> gen	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	LQ mass LQ <sup>a</sup> mass LQ <sup>a</sup> mass LQ <sup>a</sup> mass LQ <sup>a</sup> mass LQ <sup>a</sup> mass	1.8 TeV 1.7 TeV 1.49 TeV 1.24 TeV 1.24 TeV 1.25 TeV 1.26 TeV 2.0 TeV 1.96 TeV	$\begin{array}{l} \beta = 1 \\ \beta = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow br) = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow br) = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow tr) = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow br) = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow br) = 1 \\ \mathcal{B}(1, \mathbf{Q}_{j}^{*} \rightarrow br) = 1, \forall M \operatorname{coupl.} \end{array}$	2006.05872 2006.05872 2030.1294 2004.14060 2101.11582 2101.12527 ATLAS-CONF-2022-052 2330.01294	FRVZ $H \rightarrow 2\gamma_d + X$ FRVZ $H \rightarrow 4\gamma_d + X$ $H \rightarrow Z_d Z_d$	2 $\ell$ + 2 displ. vertices 2 $\mu$ -jets 2 $\mu$ -jets displaced dimuon 2 e, $\mu$ + low-EMF trackless jet	139     s lifetime       139     γ <sub>d</sub> lifetime       139     γ <sub>d</sub> lifetime       32.9     Z <sub>d</sub> lifetime	4-85 mm	0.654-939 mm 2.7-534 mm 0.21-5.2 m	$m(s) = 35 \text{ GeV}$ $m(\gamma_d) = 400 \text{ MeV}$ $m(\gamma_d) = 400 \text{ MeV}$ $m(\chi_d) = 400 \text{ MeV}$ $m(Z_d) = 40 \text{ GeV}$	2107.06092 2206.12181 2206.12181 1808.03057
$\begin{array}{c} \mathbb{V} LQ \ TT \rightarrow Zt + X \\ \mathbb{V} LQ \ BB \rightarrow Wt/Zb + X \end{array}$	2e/2µ/≥3e,µ ≥1 b, ≥1 j – 139 multi-channel 36.1		1.46 TeV 1.34 TeV	SU(2) doublet SU(2) doublet	2210.15413 1808.02343	$P_{H} \rightarrow ZZ_{d}$ $\Phi(200 \text{ GeV}) \rightarrow ss$	low-EMF trk-less jets, MS vtx			0.21-5.2 m	$m(Z_d) = 10 \text{ GeV}$ $\sigma \times \mathcal{B} = 1 \text{ pb}, m(s) = 50 \text{ GeV}$	1811.02542
VLQ $T_{5/3}T_{5/3} T_{5/3} \rightarrow Wt + J$ Q VLQ $T \rightarrow Ht/Zt$	1 e, μ ≥1 b, ≥3 j Yes 139	T <sub>5/3</sub> mass T mass	1.64 TeV 1.8 TeV	$\mathcal{B}(T_{5/3} \rightarrow Wt) = 1, c(T_{5/3}Wt) =$ SU(2) singlet, $\kappa_T = 0.5$	1 1807.11883 ATLAS-CONF-2021-040	$\Phi(600 \text{ GeV}) \rightarrow ss$	low-EMF trk-less jets, MS vtx		0.04	0.41-51.5 m	$\sigma \times \mathcal{B} = 1 \text{ pb}, m(s) = 50 \text{ GeV}$ $\sigma \times \mathcal{B} = 1 \text{ pb}, m(s) = 50 \text{ GeV}$	
$\begin{array}{c} VLQ \ Y \to Wb \\ VLQ \ B \to Hb \\ VLL \ \tau' \to Z\tau/H\tau \end{array}$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		1.85 TeV 2.0 TeV 898 GeV	$\mathcal{B}(Y \rightarrow Wb)=1, c_R(Wb)=1$ SU(2) doublet, $\kappa_B=0.3$ SU(2) doublet	1812.07343 ATLAS-CONF-2021-018 2303.05441	$\Phi(1 \text{ TeV}) \rightarrow ss$	low-EMF trk-less jets, MS vtx			06-52.4 m	$\sigma \times \mathcal{B} = 1 \text{ pb}, m(s) = 50 \text{ GeV}$ $\sigma \times \mathcal{B} = 1 \text{ pb}, m(s) = 150 \text{ GeV}$	
Excited quark $q^* \rightarrow qg$ Excited quark $q^* \rightarrow q\gamma$	- 2j - 139 1γ 1j - 36.7	q* mass q* mass	5	6.7 TeV only $u^*$ and $d^*$ , $\Lambda = m(q^*)$ i.3 TeV only $u^*$ and $d^*$ , $\Lambda = m(q^*)$	1910.08447 1709.10440	$W \to N\ell, N \to \ell\ell\nu$	displaced vtx ( $\mu\mu$ , $\mu e$ , $ee$ ) + $\mu$	139 N lifetime	0.74-42 mm		m(N) = 6 GeV, Dirac	2204.11988
Excited quark $b^* \rightarrow bg$ Excited lepton $\tau^*$	- 1 b, 1 j - 139 2 τ ≥2 j - 139		3.2 TeV 4.6		1910.08447 2303.09444	$W \to N\ell, N \to \ell\ell\nu$	displaced vtx ( $\mu\mu$ , $\mu e$ , $ee$ ) + $\mu$	139 N lifetime	3.1-33 mm		m(N)= 6 GeV, Majorana	2204.11988
Type III Seesaw LRSM Majorana $v$ Higgs triplet $H^{\pm\pm} \rightarrow W^{\pm}W^{\pm}$ Higgs triplet $H^{\pm\pm} \rightarrow \ell\ell$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	N <sub>R</sub> mass H <sup>±±</sup> mass 35	910 GeV 3.2 TeV	$m(W_R) = 4.1$ TeV, $g_L = g_R$ DY production DY production	2202.02039 1809.11105 2101.11961 2211.07505	$ \begin{array}{l} \underbrace{W}{\to} N\ell, N \to \ell\ell\nu \\ \\ W \to N\ell, N \to \ell\ell\nu \end{array} $	displaced vtx ( $\mu\mu$ , $\mu e$ , $ee$ ) + $e$ displaced vtx ( $\mu\mu$ , $\mu e$ , $ee$ ) + $e$		0.49-81 mm 0.39-51 mm		m(N)= 6 GeV, Dirac $m(N)=$ 6 GeV, Majorana	2204.11988 2204.11988
O Multi-charged particles Magnetic monopoles	$\sqrt{s} = 13 \text{ TeV}$ $\sqrt{s} = 13 \text{ TeV}$	multi-charged particle mass	1.59 TeV 2.37 TeV	$\begin{array}{l} DY \mbox{ production, }  q  = 5e \\ DY \mbox{ production, }  g  = 1g_D, \mbox{ spin} \end{array}$	ATLAS-CONF-2022-034		$\sqrt{s} = 13$ TeV $\sqrt{s} = 13$ T		0.001 0.01 0.	1 1 10	<sup>100</sup> cτ [m]	]
	partial data full data	10 <sup>-1</sup>	1	<sup>10</sup> Mass scale [Te	V]		partial data full dat	a	<u></u>			-
,	ble mass limits on new states or ph	enomena is shown.				*Only a selection of the a	available lifetime limits is	<sub>shown.</sub> 0.001	0.01 0.1	1 10	100 <b>~ [ne</b> ]	1
†Small-radius (large-radius) je	ets are denoted by the letter j (J).										au [ns	J

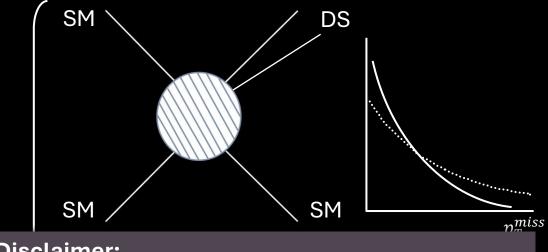


Dark Sectors at colliders





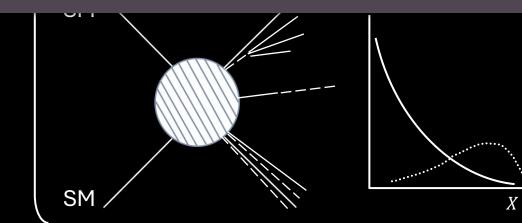




Disclaimer:

Precision measurements are also powerful probes for DSs, but focus on well-predicted SM observables which are sensitive to corrections from DS effects

Since they are experimentally different from DS searches, they will not be covered in this talk



#### MET+X searches

- DS produced recoiling against SM system
- Missing transverse energy since DS is invisible

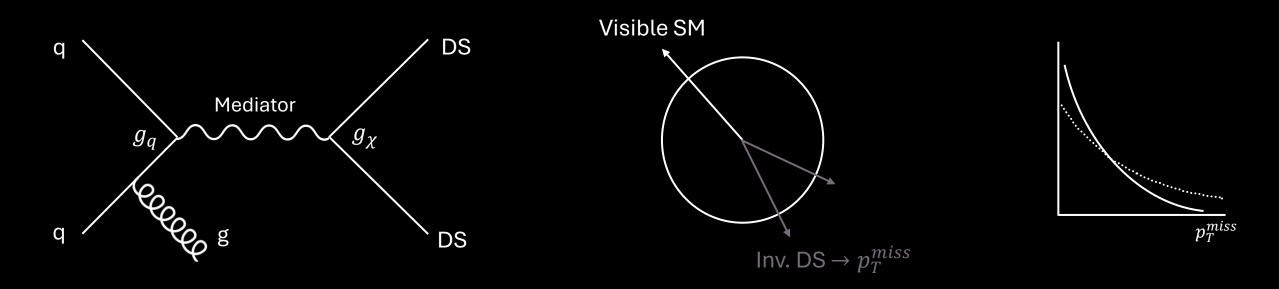
#### Portal resonances

- Known SM processes cross sections are affected by DS
- Search for bumps in mass distributions

### Unconventional signatures

- More complicated DSs can produce signatures completely different from SM (disappearing tracks, emerging jets, displaced leptons, etc.)
- New reconstructed objects often necessary

### MET+X



#### Strategy

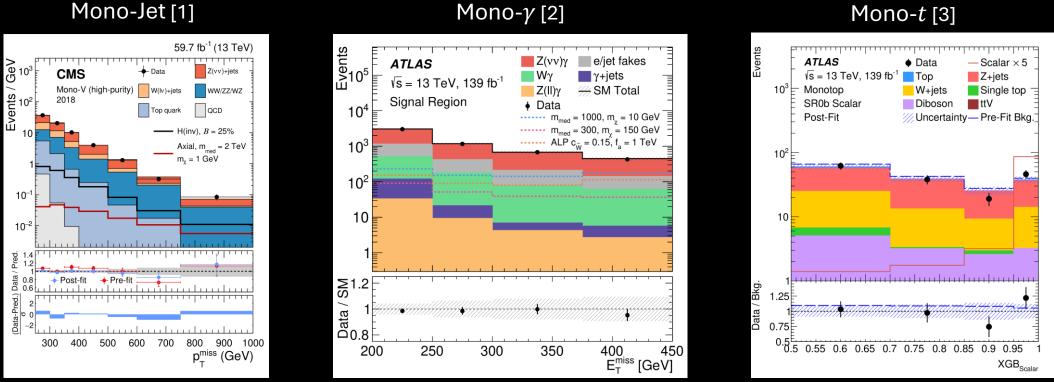
- Invisible DS particles produced via mediator that couples to SM and DS
- DS particles recoil against SM (jet, photon, V, Higgs, t/b, tt/bb, etc.)
- Since (transverse) momentum is conserved, measure missing (transverse) momentum

### Target

- Simplified DM models (e.g. WIMPs) with parameters:  $m_{med}$ ,  $m_{DM}$ ,  $g_q$ ,  $g_{\chi}$
- Higgs portals
- Any model with invisible decays! Very model independent search

# **MET+X Results**

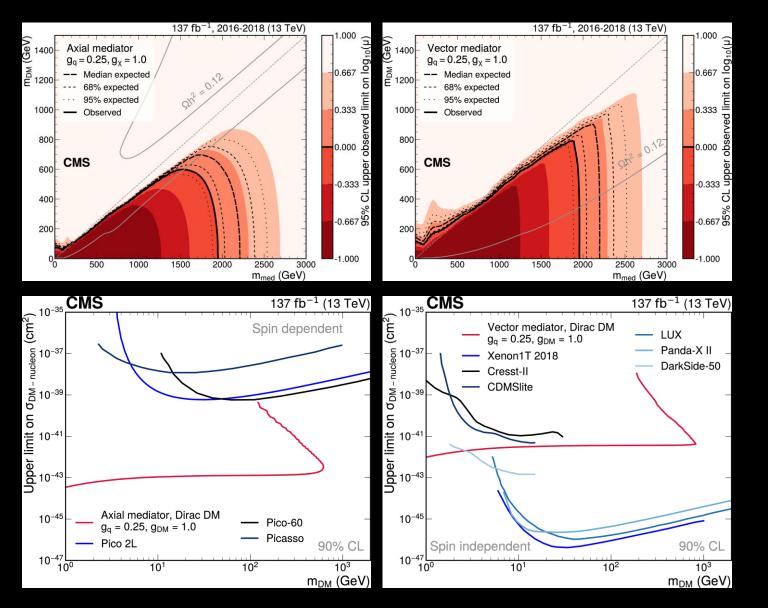
[1] - <u>JHEP 11 (2021) 153</u> [2] - JHEP 02 (2021) 226 [3] - arXiv:2402.16561



- A few representative examples of the many Run 2 MET+X results from ATLAS and CMS are shown •
- Evolution of MET algorithms to improve sensitivity (pile-up mitigation, ML, etc.)  $\bullet$
- "Control regions" in data to constrain and/or predict backgrounds  $\bullet$ 
  - Often through simultaneous binned likelihood fits with signal regions  $\bullet$

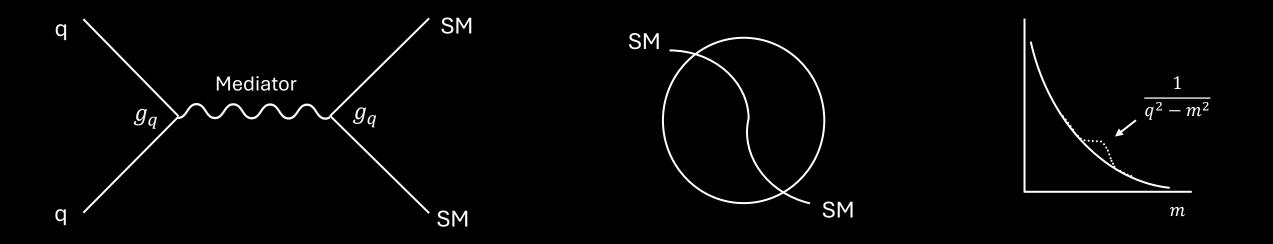
#### JHEP 11 (2021) 153

# Limits from Mono-Jet



- As an example, the same mono-jet search can be re-interpreted for many DS/DM models
  - Simplified DM models: WIMPs with vector, axial, pseudoscalars, fermion portals
  - $B(H \rightarrow inv)$
  - Leptoquarks & other more complex models
- For WIMPs, can constrain directly  $m_{DM}$  and  $m_{med}$ 
  - Can interpret these as limits on  $\sigma_{DM-nucleon}$
  - Compare with direct-detection experiments!

### **Resonance Searches**



#### Strategy

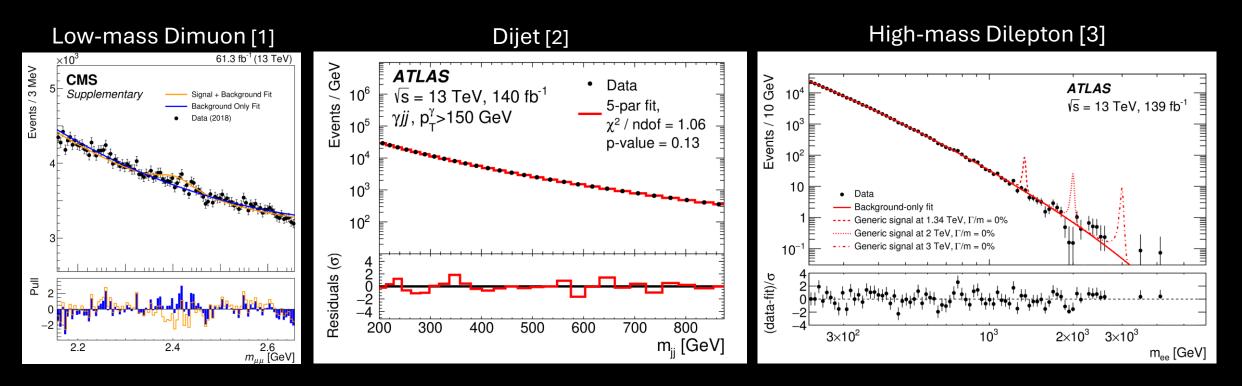
- New DS-SM mediator produced in pp collisions
- Mediator decays back to SM (instead of decaying to DS like in MET+X scenario)
- Look for Breit-Wigner resonances "bumps" in mass distributions

### Target

- Model-independent limits on  $\sigma(pp \to X)B(X \to SM SM)A$  as function of  $m_{med}$
- Similar models to MET+X, since if it can be produced via SM, it can decay back to it

## **Resonance Searches**

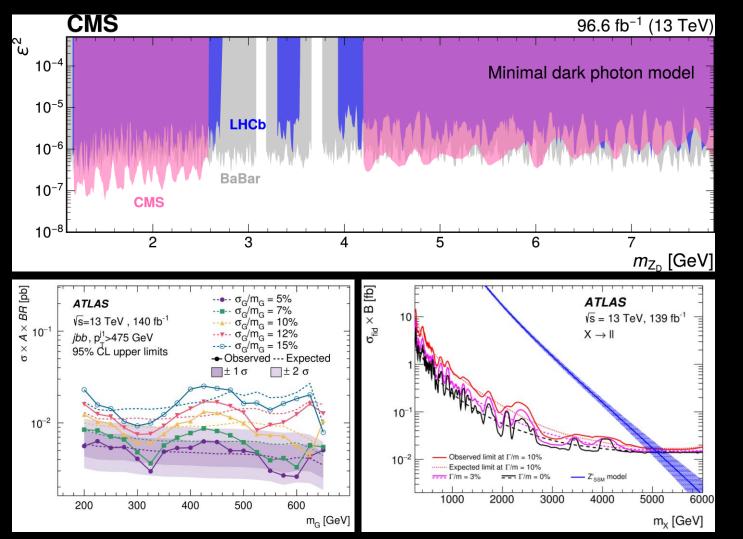
[1] - <u>JHEP 12 (2023) 070</u> [2] - <u>arXiv:2403.08547</u> [3] - <u>Phys. Lett. B 796 (2019) 68</u>



- Target high masses (~TeV) via traditional triggers
- Target low masses (~GeV) via production of another particle to trigger on (e.g. photon + 2 jets)
- Enhance sensitivity to low masses via high-rate ("scouting") triggers that select a larger fraction of signal-like events, but record less event information
- Parametrized background distributions determined from Monte Carlo, corrected in data
- "Bump-hunting": fit generic signal Breit-Wigner bumps convoluted with the detector resolution

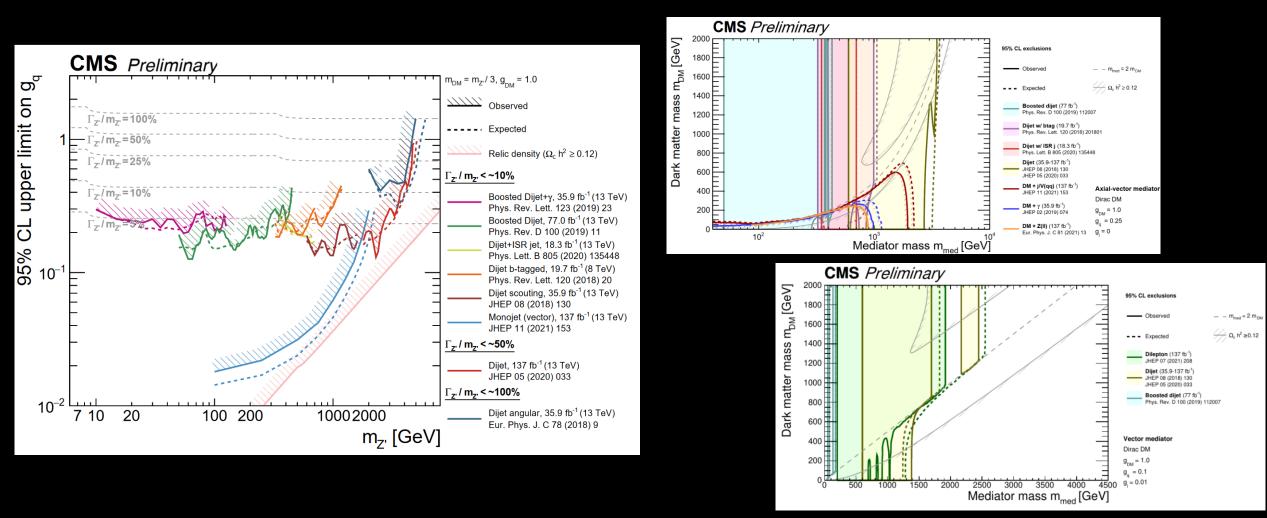
### Results from (some) Resonance Searches

<u>JHEP 12 (2023) 070</u> arXiv:2403.08547 Phys. Lett. B 796 (2019) 68



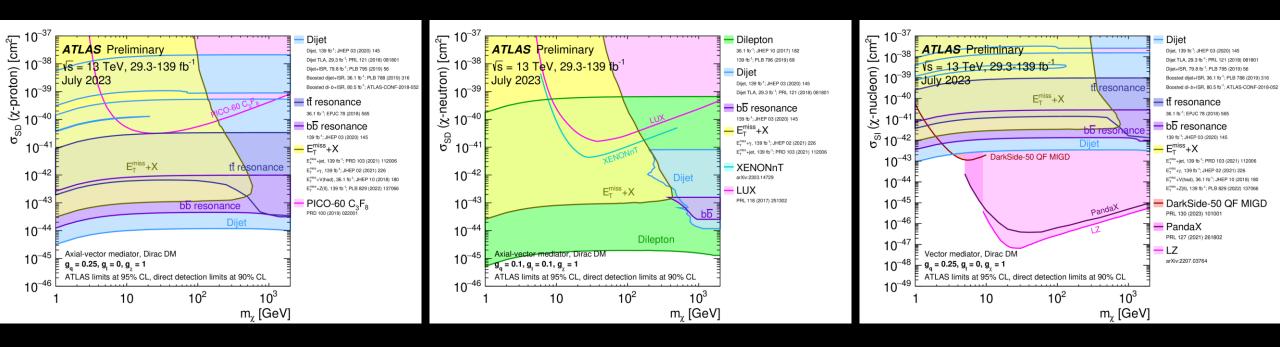
- Place limits based on the mass of the resonance and the cross section
- Dark photon model commonly used for benchmarking with other experiments, relies on mixing parameter  $\epsilon$  between the  $U(1)_D$  and SM hypercharge
- Model independent limits can be placed on simple Gaussian bumps at different m values, with different widths  $\Gamma$

### **Ground Covered**



MET+X and resonance searches have excluded large phase space of simplified DM models like WIMPs throughout Run 1 and 2 of the LHC

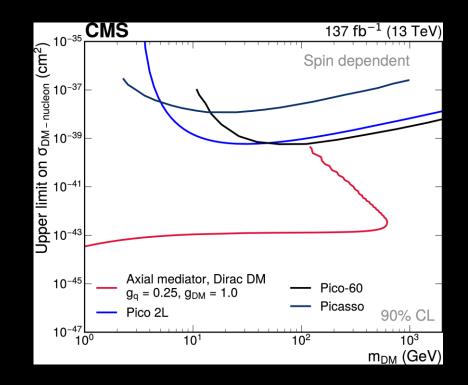
# **Colliders and Direct Detection**



- Important complementarity between colliders and direct detection experiments for simplified DM models
  - Spin dependence
  - Nature of mediator
  - Nature of dark matter particle(s)

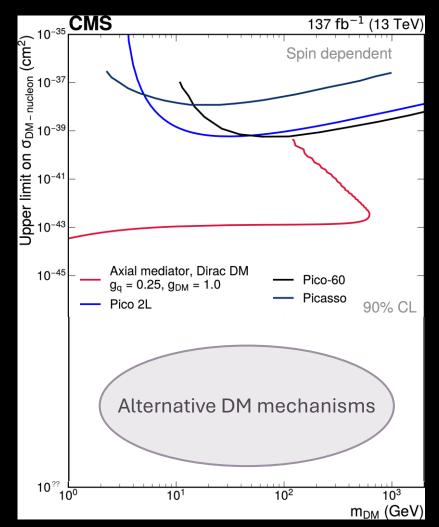
# **Unconventional Signatures**

- First-generation of searches at colliders found no convincing evidence for BSM
  - Excellent limits on simplified models have been placed
  - New ideas (scouting, ML, etc.) are still able to improve sensitivity, but we are reaching the limits of what can be done with current colliders
  - Will be re-iterated with Run 3 data (ongoing!)



# **Unconventional Signatures**

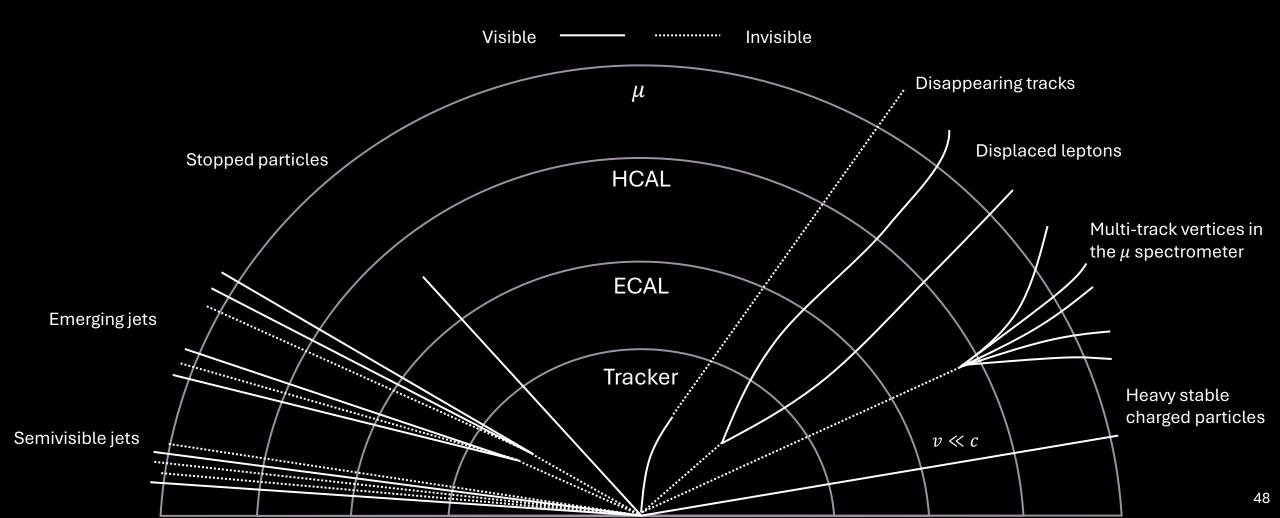
- First-generation of searches at colliders found no convincing evidence for BSM
  - Excellent limits on simplified models have been placed
  - New ideas (scouting, ML, etc.) are still able to improve sensitivity, but we are reaching the limits of what can be done with current colliders
  - Will be re-iterated with Run 3 data (ongoing!)
- More complex DS models and/or alternative DM mechanisms (non WIMP) being investigated
  - Freeze-in, inelastic DM, FIMPs, etc.
  - Complex DSs (e.g. dark QCD) could contain a stable (DM) particle as well as an unstable particles that could decay in our detectors (e.g. LLPs)
- Give rise to new types of signatures that we don't typically reconstruct at colliders
  - We would not have seen these objects at all
  - Would have evaded all previous constraints



Schema stolen from André Lessa

### **Unconventional Signatures**

Some examples of unconventional signatures that would have evaded typical reconstruction and triggers



## **Unconventional Signatures: New Approaches**

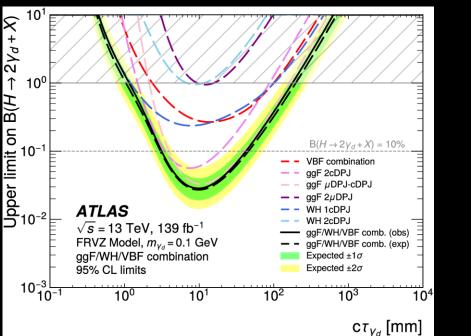
Need new approaches to reconstruct these signatures:

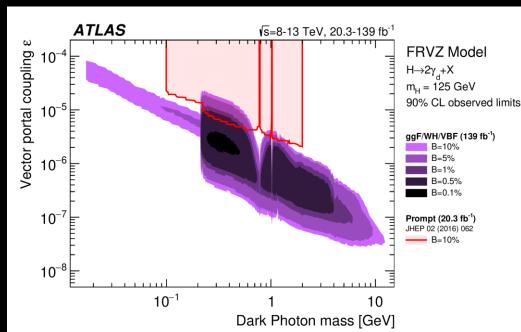
- New data streams:
  - **Scouting:** high-rate triggers, save quickly less info per event
  - **Parking**: low-rate triggers, save large amount of raw detector data to be reconstructed later
- New triggers:
  - Many dedicated new triggers to target unconventional topologies
- New offline reconstructions:
  - Looking at physics objects that may not conform to traditional, SM-like objects

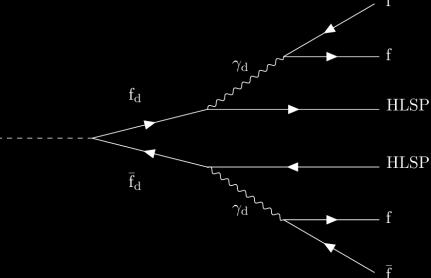
### Important part of Run 3 (2023-2025) is to leverage these new approaches

# Long Lived Dark Photons

- Displaced, collimated SM fermions reconstructed in the calorimeter or  $\mu$  spectrometer (MS)
  - Hidden lightest stable particle contributes only to  $p_T^{miss}$
- Two new triggers to target signal:
  - 3 μ's using the MS only <u>JINST 15 (2020) P09015</u>
  - 1  $\mu$  in the MS + 1  $\mu$  within  $\Delta R < 0.4$  of the first JINST 8 (2013) P07015

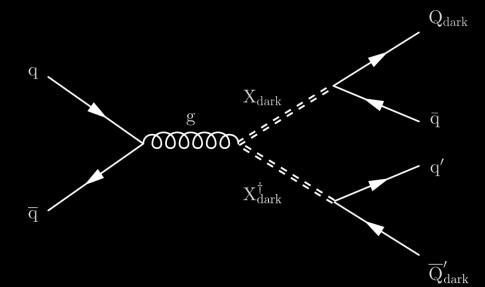


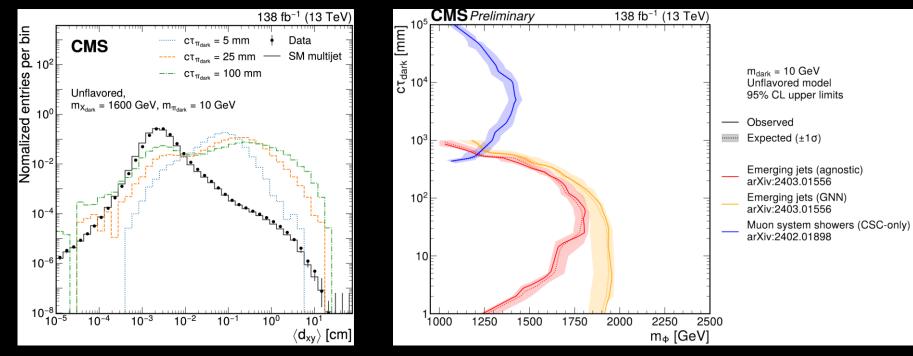




# **Emerging Jets**

- $X_{dark}$  produced, travel ~cm, decay to  $Q_{dark}$  and q
  - Two showers not associated to primary vertex
- Target jets with tracks with large displacement in the plane orthogonal to the beam  $d_{xy}$
- Alternative approach with a Graph Neural Network (GNN)





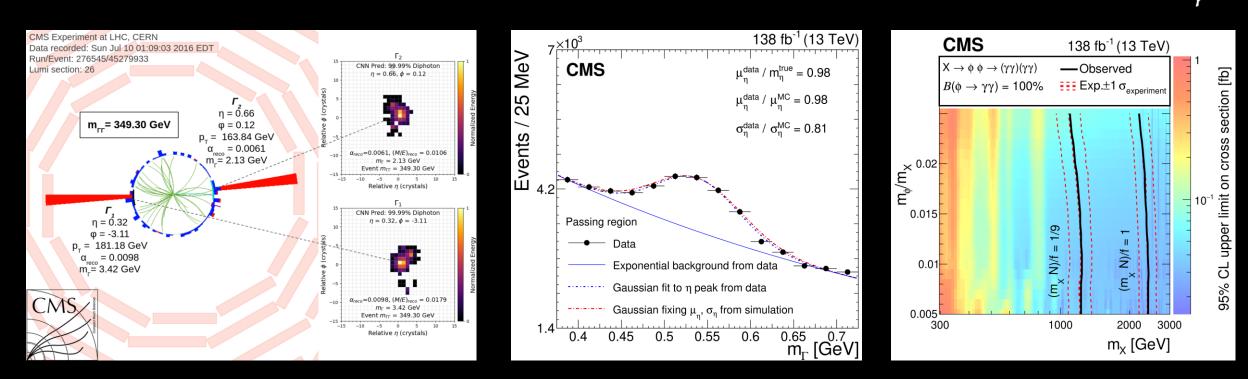
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# Merged Diphotons

- Two photons from  $\phi$  decay too merged to look like distinct photons, but not so collimated that they would look like one photon
- Dedicated CNN developed to analyze ECAL deposits
  - Distinguish single  $\gamma$ , two  $\gamma$ 's, or hadronic activity
- Second CNN to reconstruct diphoton mass
- Validate with boosted  $\pi^0$ ,  $\eta$  decays



# Conclusions

- The LHC has covered a lot of ground in DM and DS searches
- Complementary approach to indirect and direct detection experiments

### • LHC Run 3

- New physics programmes are being developed for BSM searches
  - Several talks at this conference cover them in more detail than I had time for
  - E.g. collider BSM sessions Wednesday at 3pm and today at 5pm
- New data sources, triggers, and reconstructions
- New dark sectors are being explored

### • HL-LHC

- Unprecedented luminosity will allow us to look at even rarer phenomena
- New detector technology will unlock new possibilities
  - e.g. trigger on tracks directly will massively improve many DS searches
- FCC-ee, Muon Collider?
  - Exciting new frontiers for our quest to explore BSM